Classical Fields II

1. Show that the affine distance

$$s \equiv \int_{a}^{b} \sqrt{\left|g_{\mu\nu} \frac{\mathrm{d}x'^{\mu}}{\mathrm{d}\lambda} \frac{\mathrm{d}x'^{\nu}}{\mathrm{d}\lambda}\right|} \,\mathrm{d}\lambda$$

is extremal along curves x^{τ} that satisfy

$$\frac{\mathrm{d}^2 {x'}^{\mu}}{\mathrm{d}s^2} + \Gamma'^{\mu}_{\kappa\alpha} \frac{\mathrm{d}{x'}^{\kappa}}{\mathrm{d}s} \frac{\mathrm{d}{x'}^{\alpha}}{\mathrm{d}s} = 0.$$

2. Prove that $A'_{\mu;\nu} - A'_{\nu;\mu} = A'_{\mu,\nu} - A'_{\nu,\mu}$.

3. Show that $\Gamma^{\mu}_{\mu\nu} = \frac{1}{2}g^{\mu\rho}\partial_{\nu}g_{\rho\mu}$. Show also that for any infinitesimal change $\delta \mathbf{A}$ in a matrix \mathbf{A} we have $\delta \ln(\det \mathbf{A}) = \operatorname{Tr}(\mathbf{A}^{-1}\delta \mathbf{A})$. Hence show that

$$\Gamma^{\mu}_{\mu\nu} = \partial_{\nu} \ln(\sqrt{g},) \tag{\dagger}$$

where $g \equiv |\det g|$.

4. For a weak gravitational field we write $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$, where all components of **h** are small. Show that under an infinitesimal coordinate change $\mathbf{x} \to \mathbf{x}' = \mathbf{x} + \boldsymbol{\xi}$, to first order in small quantities we have

$$h_{\mu
u}
ightarrow h'_{\mu
u} = h_{\mu
u} - \partial_{\mu}\xi_{
u} - \partial_{
u}\xi_{\mu}.$$

Explain by analogy with electromagnetism why this can be considered a gauge transfomation.

5. Show that under a gauge transformation $\mathbf{x} \to \mathbf{x}' = \mathbf{x} + \boldsymbol{\xi}$, the Christoffel symbol transforms as

$$\Gamma_{\mu\nu}^{\prime\lambda} = \frac{\partial x^{\prime\lambda}}{\partial x^{\rho}} \frac{\partial x^{\tau}}{\partial x^{\prime\mu}} \frac{\partial x^{\sigma}}{\partial x^{\prime\nu}} \Gamma_{\tau\sigma}^{\rho} - \frac{\partial x^{\rho}}{\partial x^{\prime\nu}} \frac{\partial x^{\sigma}}{\partial x^{\prime\mu}} \frac{\partial^2 x^{\prime\lambda}}{\partial x^{\rho} x^{\sigma}}$$

Hence show that we can always choose a gauge in which the harmonic gauge condition

$$\Gamma^{\lambda} \equiv g^{\mu\nu} \Gamma^{\lambda}_{\mu\nu} = 0$$

is satisfied.

Use the result (†) above to show that the harmonic gauge condition can be writen

$$0 = \partial_{\kappa} (\sqrt{g} \, g^{\lambda \kappa}) \,.$$

Hence show that when the harmonic gauge condition holds, the covariant d'Alembertian

$$\Box \phi \equiv \nabla_{\kappa} (g^{\kappa \lambda} \nabla_{\lambda} \phi)$$

is simply

$$\Box \phi = g^{\lambda \kappa} \frac{\partial^2 \phi}{\partial x^\lambda \partial x^\kappa}.$$

Thus when the harmonic gauge condition is satisfied, each coordinate is a harmonic function: $\Box x^{\alpha} = 0$.

6. Show that

$$g_{\lambda\sigma}\partial_{\kappa}g^{\sigma\rho} = -g^{\sigma\rho} \big(\Gamma^{\eta}_{\kappa\lambda}g_{\eta\sigma} + \Gamma^{\eta}_{\kappa\sigma}g_{\eta\lambda}\big).$$

Hence, or otherwise, show that the curvature tensor can be written

$$R_{\lambda\mu\nu\kappa} = \frac{1}{2} \left(\frac{\partial^2 g_{\lambda\nu}}{\partial x^{\kappa} \partial x^{\mu}} - \frac{\partial^2 g_{\mu\nu}}{\partial x^{\kappa} \partial x^{\lambda}} - \frac{\partial^2 g_{\lambda\kappa}}{\partial x^{\nu} \partial x^{\mu}} + \frac{\partial^2 g_{\mu\kappa}}{\partial x^{\nu} \partial x^{\lambda}} \right) + g_{\eta\sigma} \left(\Gamma^{\eta}_{\nu\lambda} \Gamma^{\sigma}_{\mu\kappa} - \Gamma^{\eta}_{\kappa\lambda} \Gamma^{\sigma}_{\mu\nu} \right).$$

Deduce that $R_{\lambda\mu\nu\kappa} + R_{\lambda\kappa\mu\nu} + R_{\lambda\nu\kappa\mu} = 0$. Hence show that $R_{\lambda\mu\nu\kappa}$ has 20 independent indices.

7. Show that for a weak field $(\mathbf{g} = \boldsymbol{\eta} + \mathbf{h})$, to first order in \mathbf{h} the Ricci tensor is

$$R_{\alpha\beta} = \frac{1}{2}\partial_{\beta}\partial_{\alpha}h - \frac{1}{2}\partial^{\nu}(\partial_{\beta}h_{\nu\alpha} + \partial_{\alpha}h_{\nu\beta} - \partial_{\nu}h_{\alpha\beta}),$$

where $h = h_{\alpha}^{\alpha}$. Show further that to first order the harmonic gauge condition reads

$$0 = 2\partial^{\mu}h_{\lambda\mu} - \partial_{\lambda}h.$$

Hence show that, in the harmonic gauge and to first order, Einstein's equations, $R_{\alpha\beta} - \frac{1}{2}Rg_{\alpha\beta} = -(8\pi G/c^4)T_{\alpha\beta}$, read

$$\Box \overline{h}_{\alpha\beta} = -\frac{16\pi G}{c^4} T_{\alpha\beta},\tag{\ddagger}$$

where $\overline{h}_{\alpha\beta} = h_{\alpha\beta} - \frac{1}{2}h\eta_{\alpha\beta}$.

Write down an analogous equation of electromagnetism and explain the physical significance of this result.

8. Use equation (‡) above to show that in harmonic coordinates the metric associated with a weak gravitational field that is caused by a static distribution of rest-mass takes the form

$$ds^{2} = -\left(1 + 2\frac{\Phi}{c^{2}}\right)c^{2}dt^{2} + \left(1 - 2\frac{\Phi}{c^{2}}\right)(dx^{2} + dy^{2} + dz^{2}),$$

where Φ should be related to T_{00} .